

# Domain wall flow equations and flux compactifications

*Paul Smyth @ Zurich Strings Workshop '09*



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based on: P.S. & S. Vaula 0905.1334 [hep-th]



# Motivation

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- ◆ Flux compactifications and their 4D low energy effective actions have seen much attention recently.

*see also the talks of Martucci, Quevedo, Roest, Wrase...*

- ◆ It has long been appreciated that the resulting 4D theories are gauged supergravities with can have electric and magnetic charges Polchinski & Strominger, Michelson '95

*recall De Rydt's talk*

*Fluxes* → *Curvature/Torsion* → *Scalar Potential in 4d*

# Aim of our work

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- ◆ Minkowski<sub>4</sub> and AdS<sub>4</sub> vacua have been studied in much detail Graña, Minasian, Petrini & Tomasiello; Lust & Tsimpis...
- ◆ Domain walls are ubiquitous vacua in gauged supergravity, yet rarely considered, so why not look at them too? Mayer & Mohaupt, Louis et al



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- ◆ Minkowski<sub>4</sub> and AdS<sub>4</sub> vacua have been studied in much detail Graña, Minasian, Petrini & Tomasiello; Lust & Tsimpis...
- ◆ Domain walls are ubiquitous vacua in gauged supergravity, yet rarely considered, so why not look at them too? Mayer & Mohaupt, Louis et al
- ◆ Our aim is to find the general set equations describing these vacua, which might allow us to find new solutions.
- ◆ We test the proposal for flux compactifications given by Graña et al using the language of  $SU(3) \times SU(3)$  structures ~ a handy way to deal with manifolds with torsion.

# Overview

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*How do we go about doing this?*



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Domain wall  
in 4D  $N=2$  theory

$$\mathbb{R}^{1,3}$$

Lifted ansatz  
in 10D *Type II* theory

$$\mathbb{R}^{1,3} \times Y$$

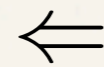
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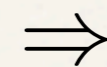
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Flow equations  
for scalar fields



SUSY variations



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Conditions on  
structures  $\Phi_{\pm}$  on  
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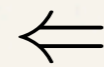
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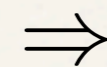
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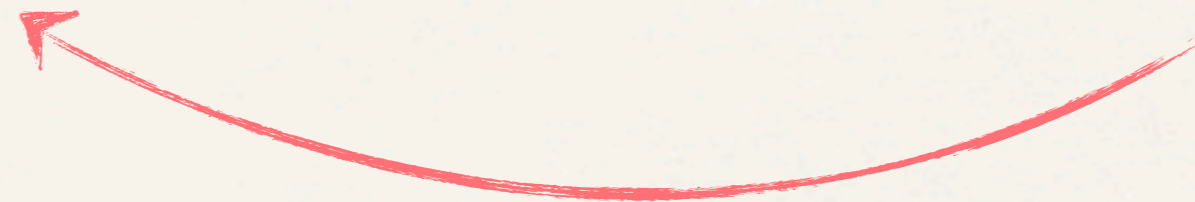
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*Graña-Louis-Waldram*  
 *$SU(3) \times SU(3)$  structure dictionary*



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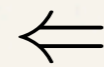
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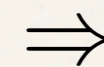
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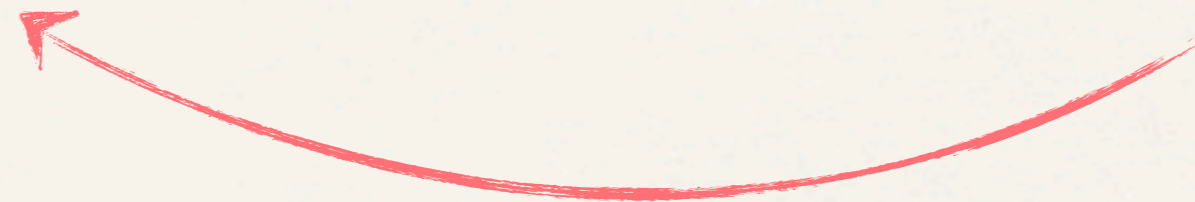
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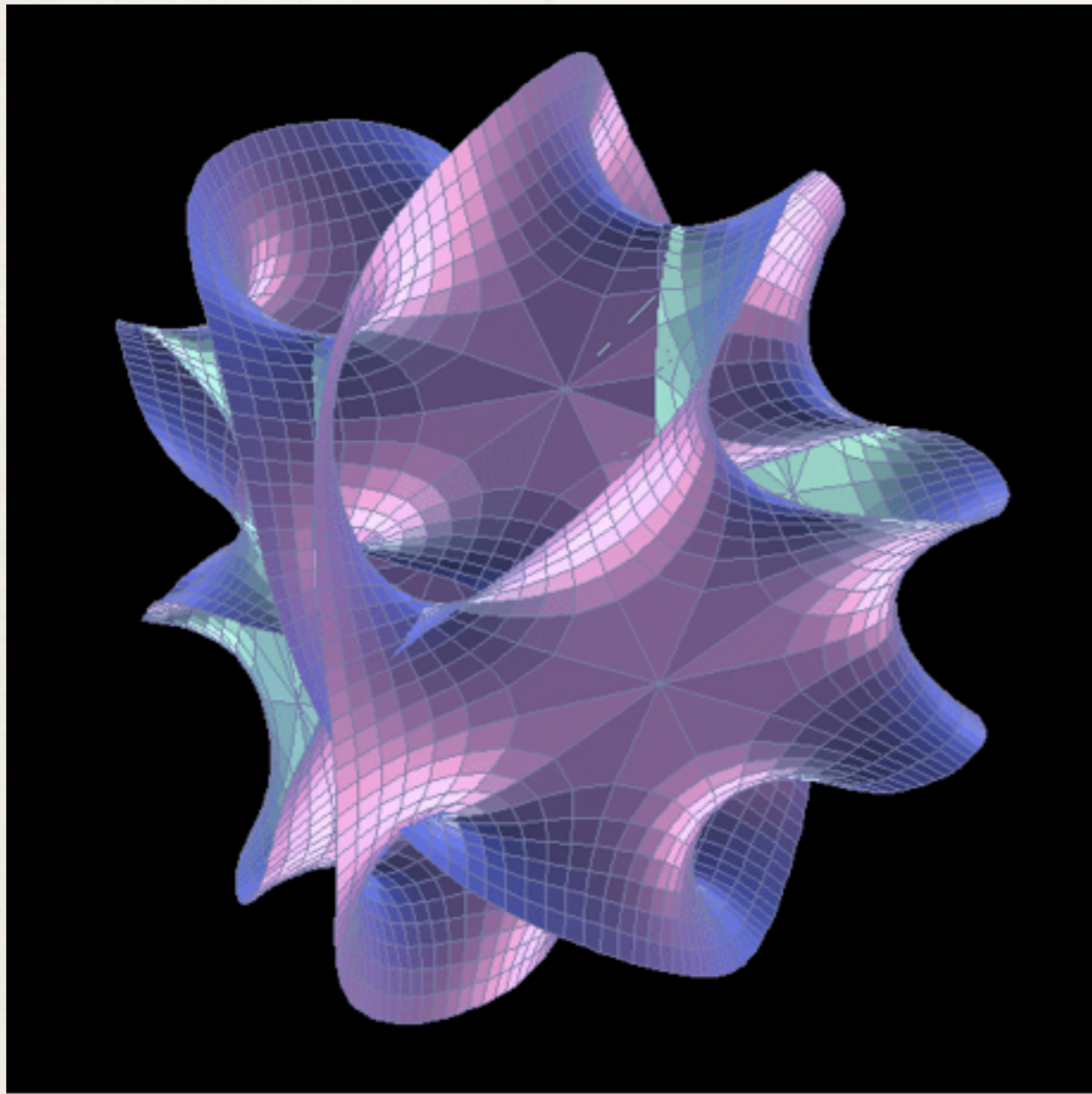
*Graña-Louis-Waldram*  
 *$SU(3) \times SU(3)$  structure dictionary*

*Test the dictionary - do the results for vacua agree?*



# From $SU(3)$ holonomy to $SU(3)$ structure

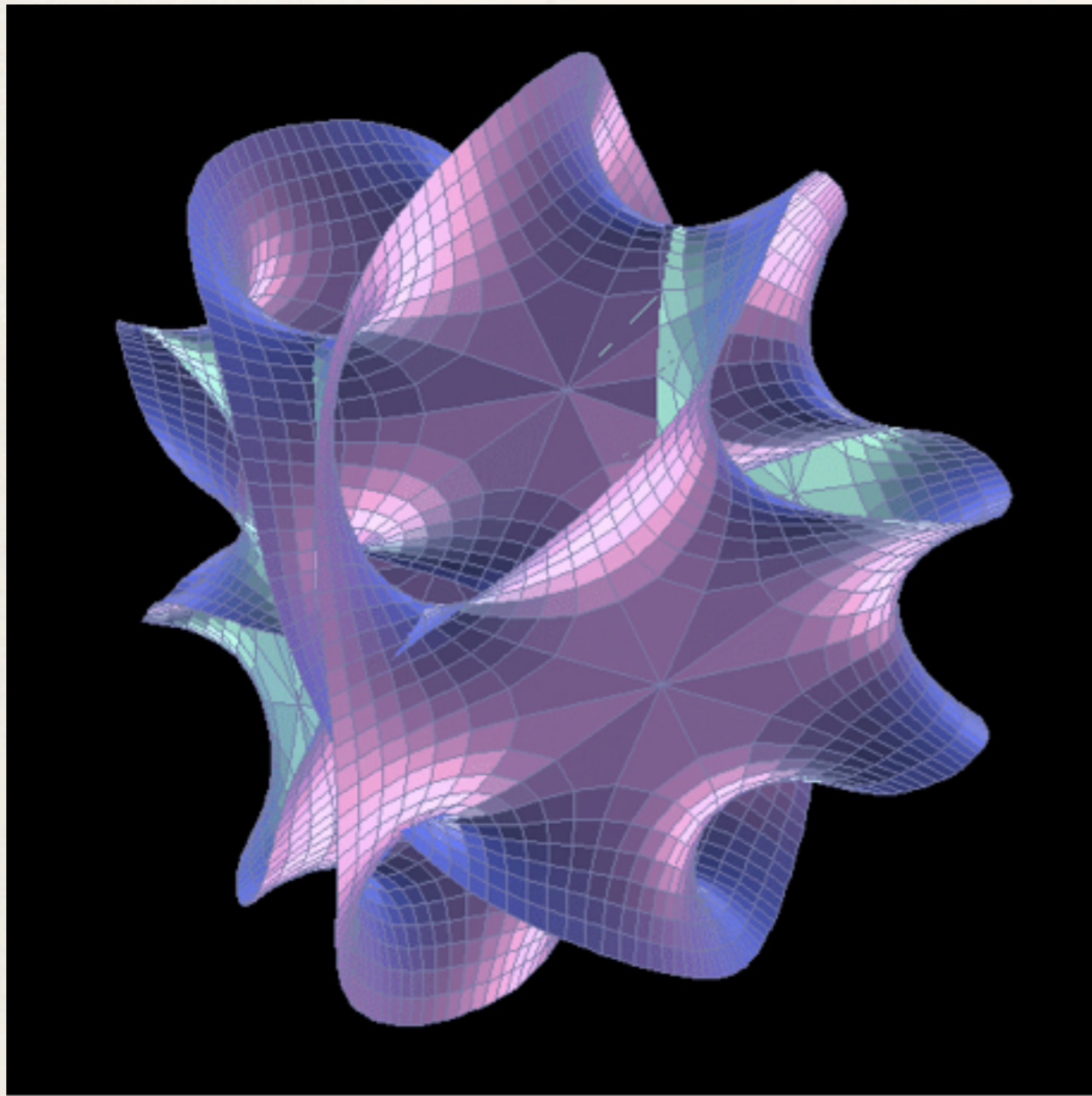
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# From $SU(3)$ holonomy to $SU(3)$ structure

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*Thanks to Jeff Bryant @ Wolfram Research and calabi yau*



# Domain Walls from 10D

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We consider an ansatz for a D=4 supersymmetric domain wall in a 10D bosonic warped background of Type II supergravity:

$$ds^2 = e^{2A(y,r)} \left( e^{2V(r)} \eta_{\alpha\beta} dx^\alpha dx^\beta + e^{2G(r)} dr^2 \right) + g_{mn}(r,y) dy^m dy^n ,$$

$$F_n^{(10)} = \hat{F}(y)_n + \text{vol}_4 \wedge \tilde{F}(y)_{n-4} , \quad H^{(10)} = H_3(y) , \quad \phi = \phi(r,y)$$

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*Use the SUSY variations to find the conditions for domain wall vacua*



# A first look at the SUSY variations

---

$$\delta\psi_M = (D_M \pm \frac{1}{4}H_M\mathcal{P})\epsilon + \frac{e^\phi}{16} \sum_n \hat{F}_{2n} \Gamma_M \mathcal{P}_n \epsilon$$

For such a symmetric configuration the SUSY variations are considerably simplified. To proceed, we still need some reasonable assumption - no worldvolume dependence:

$$\epsilon(x^\alpha, r, y^m) = \epsilon(r, y^m)$$

Radial dependence plays an important role, so we can initial focus on this component of the gravitino variation  $\delta\psi_r$  and  $\delta\psi_\alpha$ :

$$\delta\psi_r \Rightarrow \epsilon(r, y^m) = e^{\frac{1}{2}(A(r)+V(r))} \epsilon_0(y^m)$$



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# 10D Killing Spinors $\epsilon_0(y^m)$

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We take the standard ansatz for the 10D Killing spinors:

$$\epsilon_0^1(y^m) = \epsilon_+ \otimes \eta_+^{(1)}(y^m) + \epsilon_- \otimes \eta_-^{(1)}(y^m)$$

$$\epsilon_0^2(y^m) = \epsilon_+ \otimes \eta_-^{(2)}(y^m) + \epsilon_- \otimes \eta_+^{(2)}(y^m)$$

where the subscripts label chirality,  $\epsilon$  is a 4D spinor and  $\eta^{(1)}$ ,  $\eta^{(2)}$  are 6D spinors on  $Y$ .

Also, we will need an ansatz for the DW projection condition:

$$\gamma_{\underline{r}} \epsilon_+ = i\alpha \epsilon_-$$

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# Flux compactifications

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*Now follow the story in Green, Schwarz and Witten 2:*

$$\eta \rightarrow \eta^{(1)} \ \& \ \eta^{(2)}$$

The spinors  $\eta^{(1)}$  &  $\eta^{(2)}$  **each** locally define an SU(3) structure on  $Y$ .

Globally, the structure can be SU(3)  $\eta^{(1)} = \eta^{(2)}$  or SU(2)  $\eta^{(1)} \neq \eta^{(2)}$ .



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$\delta\psi_m$  &  $\delta\lambda$  then give conditions on the spinors on  $Y$ , or on the associated structures  $\sim dJ=0$  and  $d\Omega=0$  in the Calabi-Yau case.



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Internal fluxes  $\Rightarrow R_Y \neq 0$  and  $Y$  is a manifold with **torsion**:  $\nabla_M \varepsilon \neq 0$



Generalised geometry provides a compact description of manifolds with torsion by considering structures on  $T_Y \oplus T^*_Y$ :

$\eta^{(1)}$  &  $\eta^{(2)}$  then define an  **$SU(3) \times SU(3)$  structure** on  $T_Y \oplus T^*_Y$ .



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$\eta^{(1)}$  &  $\eta^{(2)}$  then define an  **$SU(3) \times SU(3)$  structure** on  $T_Y \oplus T^*_Y$ .

Use the tensor product to define *pure spinors* on  $T_Y \oplus T^*_Y$ :

$$\Phi_{\pm} = \eta_{+}^{(1)} \otimes \eta_{\pm}^{(2)\dagger} \quad \leftrightarrow \quad \text{Fierz id.} \equiv \text{Sums of even/odd forms}$$

purity means  $\Phi_{\pm}$  is annihilated by a 6D subspace of Clifford(6,6).

*c.f. the definition in Boels talk.*



# Pure spinor supersymmetry conditions

---

Using the tools of AdS<sub>4</sub> compactifications, we can take the conditions on the  $\eta$ 's for supersymmetric domain walls found from  $\delta\psi_m$  &  $\delta\lambda$  and rewrite them in terms of pure spinors  $\Phi_{\pm}$ :

$$\begin{aligned}d_H [e^{2A-\phi} \text{Im}\Phi_-] &= 0 , \\d_H [e^{4A-\phi} \text{Re}\Phi_-] &= e^{4A} \tilde{F} - e^{-3V-G} \text{Im} (\alpha^* \partial_r [e^{3A+3V-\phi} \Phi_+]) , \\d_H [e^{3A-\phi} \text{Im} (\alpha^* \Phi_+)] &= 0 , \\d_H [e^{3A-\phi} \text{Re} (\alpha^* \Phi_+)] &= e^{-2V-G} \partial_r \text{Im} [e^{2A+2V-\phi} \Phi_-] ,\end{aligned}$$

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Domain wall flow equations



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Radial flow terms reduce to the *cosmological constant* in the AdS limit



## Example: Domain walls and *Half-flat* SU(3) structure

Pure spinors

$$\Phi_+ = e^{iJ} , \quad \Phi_- = \Omega , \quad \left( \eta^{(1)} = \eta^{(2)} \right)$$



## Example: Domain walls and *Half-flat* SU(3) structure

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$$d\text{Re}\Phi_+ \equiv -d(J \wedge J) = 0 ,$$

Differential  
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$$d\text{Im}\Phi_+ \equiv dJ = -\partial_r(\text{Im } \Omega),$$

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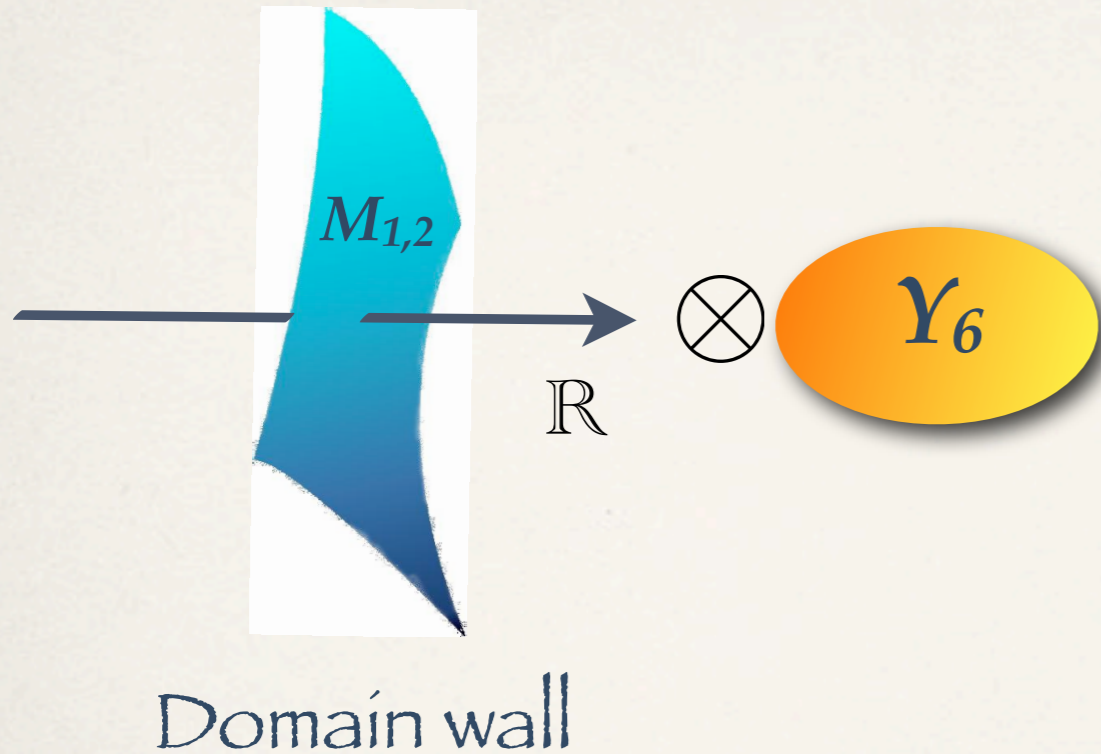
Defining 3 & 4 form in 7 dimensions (i.e.  $Y + dr$ ) we find,

$$\rho = dr \wedge J + \text{Im}(\Omega) \quad , \quad \star \rho = dr \wedge \text{Re}(\Omega) + J \wedge J$$

SU(3) structure  $\hookrightarrow$  G<sub>2</sub> holonomy :  $d_7\rho = 0 = d_7(\star\rho)$

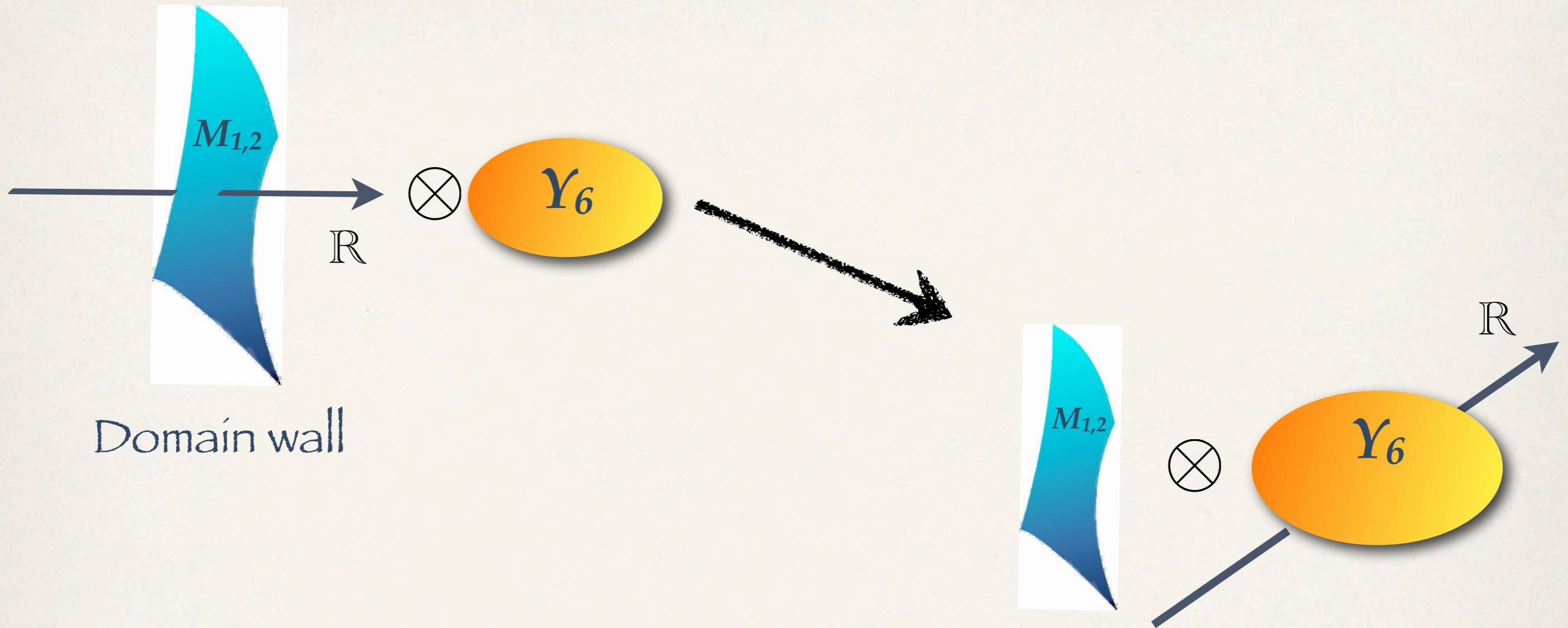


# From $SU(3)$ structure to $G_2$ holonomy for domain walls



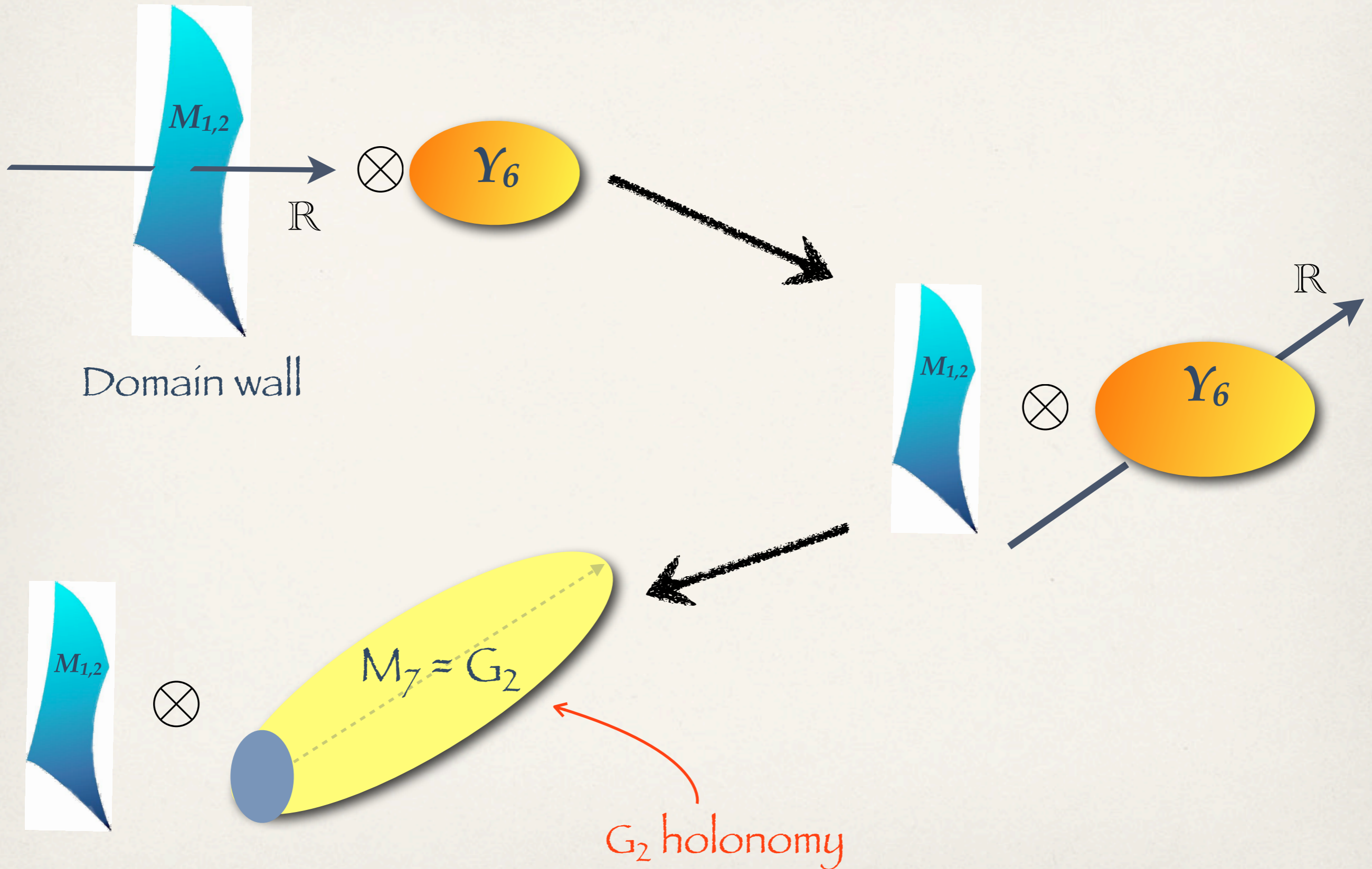


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Our 10D result for the  $SU(3) \times SU(3)$  structure case:

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which are *generalised Hitchin flow equations*.

They describe the embedding of the  $SU(3) \times SU(3)$  structure manifold into an *generalised  $G_2$  manifold* i.e.  $G_2 \times G_2$  structure. Jeschek & Witt '05



# A quick recap

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- ◆ We used the 10D SUSY variations to derive a set of equations describing a domain wall in 4D.
- ◆ This approach obviously does not require any truncation of fields, therefore it will give a useful tool to compare with the truncated 4d results.
- ◆ The resulting 'flow equations' describe the embedding of a  $SU(3) \times SU(3)$  structure manifold into a generalised  $G_2$  manifold.



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*Do they agree with the 4D gauged supergravity equations?*



# The 4D story

---

*Now forget everything and start again...*

We will consider  $N=2$  supergravity in 4d, which can contain vector- and hyper-multiplets with scalar fields  $t^i$  and  $q^u$  respectively.

For particular gaugings, the  $N=2$  theory is believed to match the proposed reduction of Type II supergravity on a  $SU(3) \times SU(3)$  structure manifold. Graña, Louis & Waldram '05-'06

We will use the GLW dictionary between 4D and 10D fields to compare the compare the flow equations.



The supersymmetry variations are:

$$\begin{aligned}
 \text{Gravitino:} \quad \delta\psi_{\mu A} &= D_{\mu}\varepsilon_A + iS_{AB}\gamma_{\mu}\varepsilon^B = 0 , \\
 \text{Gaugino:} \quad \delta\lambda^{iA} &= i\partial_{\mu}t^i\gamma^{\mu}\varepsilon^A + W^{iAB}\varepsilon_B = 0 , \\
 \text{Hyperino:} \quad \delta\zeta_{\hat{\alpha}} &= iP_{uA\hat{\alpha}}\partial_{\mu}q^u\gamma^{\mu}\varepsilon^A + N_{\hat{\alpha}}^A\varepsilon_A = 0 ,
 \end{aligned}$$

where

$$\begin{aligned}
 S_{AB} &= \frac{i}{2}\vec{\sigma}_{AB} \cdot \vec{W} , \\
 W^{iAB} &= i\vec{\sigma}^{AB} g^{i\bar{j}}\nabla_{\bar{j}}\vec{W} , \\
 P_{(A}{}^{v\hat{\alpha}}N_{B)\hat{\alpha}} &= i\vec{\sigma}^{AB} h^{vu}\nabla_u\vec{W} ,
 \end{aligned}$$

$g^{i\bar{j}}$  &  $h^{vu}$  are the vector- and hyperscalar  $\sigma$ -model metrics,  $P_{uA\hat{\alpha}}$  is the quaternionic vielbein and  $\vec{W}$  is a triplet of superpotentials.



# Domain wall ansatz again...

---

Now we play the same game, plugging a domain wall ansatz into the SUSY variations:

$$ds_4^2 = e^{2U(r)} \eta_{\alpha\beta} dx^\alpha dx^\beta + e^{-2pU(r)} dr^2 ,$$

$$t^i(x^\alpha, r) = t^i(r) \qquad q^u(x^\alpha, r) = q^u(r)$$

Using some standard tricks, we find:

$$\partial_r t^i = \mp e^{-pU} g^{i\bar{j}} \nabla_{\bar{j}} \bar{W} ,$$

$$\partial_r q^u = \mp e^{-pU} h^{uv} \partial_v \bar{W} ,$$

$$U' = \pm e^{-pU} W ,$$

*How can we compare this to the 10D result?*



We need to express the 4D quantities in terms of  $\Phi_{\pm}$  and  $\tilde{F}$ :

hyper:  $q^u \begin{cases} \nearrow z^a \rightarrow \Phi_-^0 \\ \searrow \xi_{RR}^A \rightarrow \tilde{F} \end{cases}$

vector:  $t^i \rightarrow \Phi_+^0$

$$\Phi_{\pm}^0 \equiv \Phi_{\pm} \Big|_{\mathcal{U}}$$



Truncated, finite-dim special Kähler subspace  $\mathcal{U}$  of pure spinors

Similarly, there is an expression for the superpotential  $\vec{W}$  in terms of  $\tilde{F}$  and  $d_H \Phi_{\pm}^0$  which we can apply on the RHS of our flow equations.



# Flow equations for domain walls in 4D

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After some work we find more complicated expressions,

$$d_H \text{Re}(e^{-\phi} h \Phi_+^0) = e^{-\varphi + pU - 2(U + \varphi)} \partial_r \text{Im}(e^{2(U + \varphi)} e^{-\phi} \Phi_-^0) ,$$

$$d_H \text{Re}(e^{-\phi} \Phi_-^0) = \tilde{F} - e^{-\varphi + pU - (U + \varphi)} \partial_r \text{Im}(e^{U + \varphi} e^{-\phi} h \Phi_+^0) ,$$

$$d_H \text{Im}(e^{-\phi} \Phi_-^0) = 0 ,$$

$$d_H \text{Im}(e^{-\phi} h \Phi_+^0) = 0 .$$

where we have also made use of the relation between the 4D and 10D dilatons,

$$\phi \propto \varphi - \frac{1}{2} K_+$$

*Now we can try to match metrics, 4D spinors, etc*



# Comparing with our result from 10D

---

- ◆ *Warp factor*  $A(y, r) = 0$  - KK reduction is better understood.
- ◆ Then we need to take care of the 4d spinors and match the metric.



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*The results agree*

- ◆ We find a precise agreement between the (truncated) 10D result and the pure 4D supergravity result.
- ◆ Examples? From the 4D perspective, we were able to find domain walls with more general NS charges (R.-R.=0).



# Conclusions

---

*What is this really good for?*

- ◆ Our results provide a non-trivial check of the generalised compactification procedure proposed by GLW.
- ◆ The SUSY variations for domain wall vacua produce *generalised Hitchin flow equations*, describing the embedding of a  $SU(3) \times SU(3)$  structure manifold into a generalised  $G_2$  manifold.



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*Some further work*

- \* More general examples and the relation to the  $N=1$  story.
- \* What about other vacua? (PS, J. Louis and H. Triendl - coming soon)