# Higher-Spin Theory in the Local Frame and Holography

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## Plan

- Motivation
- Free HS fields
- HS symmetries versus locality
- Unfolded dynamics
- Perturbative analysis and resolution operator for nonlinear HS equations
- Local HS cubic interactions and holography
- Functional class for HS fields
- Conclusion

## **Challenge: Quantum Gravity**

Conjecture that the regime of ultra high (transPlanckian) energies exhibits some high symmetries

The idea of HS gauge theory is to understand what kind of higher symmetries can be introduced in relativistic theory

It was argued (Weinberg, Coleman-Mandula, Weinberg-Witten) that HS symmetries cannot be realized in a nontrivial local field theory in Minkowski space

In 70th it was shown by Aragone and Deser that HS gauge symmetries are incompatible with GR if expanding around Minkowski space

**Green light:** (*A*)*dS* **background with**  $\Lambda \neq 0$  Fradkin, MV, 1987 **In agreement with no-go statements the limit**  $\Lambda \rightarrow 0$  **is singular** 

### **HS** Theories and String Theory

**HS theories:**  $\Lambda \neq 0$ , m = 0

symmetric fields  $s = 0, 1, 2, \dots \infty$ 

String Theory:  $\Lambda = 0$ ,  $m \neq 0$  except for a few zero modes mixed symmetry fields  $\overrightarrow{s} = 0, 1, 2, ... \infty$ 

String theory has much larger spectrum: HS Theory: first Regge trajectory

Pattern of HS gauge theory is determined by HS symmetry

What is a string-like extension of a global HS symmetry underlying a string-like extension of HS theory?

MV 2012, Gaberdiel and Gopakumar 2014-2017

• String Theory as spontaneously broken HS theory?! (s > 2, m > 0)

## **Fronsdal Fields**

**Fronsdal fields** 1978 **All** m = 0 **HS fields are gauge fields**   $\varphi_{n_1...n_s}$  **is a rank** *s* **symmetric tensor obeying**  $\varphi^k_k{}^m_{mn_5...n_s} = 0$ **Gauge transformation:** 

$$\delta \varphi_{n_1 \dots n_s} = \partial_{(n_1} \varepsilon_{n_2 \dots n_s)}, \qquad \varepsilon^m_{m n_3 \dots n_{s-1}} = 0$$

Fronsdal action  $S(\varphi)$  implies field equations:  $R_{n_1...n_s}(x) = 0$  $R_{n_1...n_s}(x)$ : Ricci-like tensor

 $R_{n_1...n_s}(x) = \Box \varphi_{n_1...n_s}(x) - s \partial_{(n_1} \partial^m \varphi_{n_2...n_sm}(x) + \frac{s(s-1)}{2} \partial_{(n_1} \partial_{n_2} \varphi^m_{n_3...n_sm}(x)$ Main principle fixing the structure of HS theory: HS gauge symmetry

# **HS Symmetries Versus Riemann Geometry**

HS symmetries do not commute with space-time symmetries

$$[T^n, T^{HS}] = T^{HS}, \qquad [T^{nm}, T^{HS}] = T^{HS}$$

HS transformations map gravitational fields (metric) to HS fields

 $\delta_{HS}\varphi_{nm}\sim\varphi_{HS}$ 

#### **Consequence:**

Riemann geometry is not appropriate for HS theory:

concept of local event may become illusive!

Cartan formalism of differential forms preserves coordinate independence without metric.

# **Unfolded Dynamics**

#### **First-order form of differential equations**

$$\dot{q}^i(t) = \varphi^i(q(t))$$
 initial values:  $q^i(t_0)$ 

Unfolded dynamics: multidimensional generalization

$$\frac{\partial}{\partial t} \to \mathsf{d} \,, \qquad q^{i}(t) \to W^{\Omega}(x) = dx^{n_{1}} \wedge \ldots \wedge dx^{n_{p}} W^{\Omega}_{n_{1} \ldots n_{p}}(x)$$
$$\mathsf{d} W^{\Omega}(x) = G^{\Omega}(W(x)) \,, \qquad \mathsf{d} = dx^{n} \partial_{n}$$

 $G^{\Omega}(W)$ : function of "supercoordinates"  $W^{\Omega}$ 

$$G^{\Omega}(W) = \sum_{n=1}^{\infty} f^{\Omega} \Phi_{1} \dots \Phi_{n} W^{\Phi_{1}} \wedge \dots \wedge W^{\Phi_{n}}$$

**Covariant first-order differential equations** 

#### d > 1: Compatibility conditions

$$G^{\Phi}(W) \wedge \frac{\partial G^{\Omega}(W)}{\partial W^{\Phi}} \equiv 0$$

## **Properties**

- General applicability
- Manifest (HS) gauge invariance under the gauge transformation

$$\delta W^{\Omega} = \mathrm{d}\varepsilon^{\Omega} + \varepsilon^{\Phi} \frac{\partial G^{\Omega}(W)}{\partial W^{\Phi}},$$

gauge parameter  $\varepsilon^{\Omega}(x)$  is a  $(p_{\Omega} - 1)$ -form.

- Invariance under diffeomorphisms Exterior algebra formalism
- Interactions: nonlinear deformation of  $G^{\Omega}(W)$ Independence of ambient space-time: geometry is encoded by  $G^{\Omega}(W)$ Key observation: unfolded equation makes sense in any space-time

$$dW^{\Omega}(x) = G^{\Omega}(W(x)), \quad x \to X = (x, z), \quad d_x \to d_X = d_x + d_z, \quad d_z = dz^u \frac{\partial}{\partial z^u}$$

*X*-dependence is reconstructed in terms of fields  $W^{\Omega}(X_0) = W^{\Omega}(x_0, z_0)$ at any  $X_0$ . To take  $W^{\Omega}(x_0, z_0)$  in space  $M_X$  with coordinates  $X_0$  is the same as to take  $W^{\Omega}(x_0)$  in the space  $M_x \in M_X$  with coordinates x

#### Frame-Like HS Fields

 $\begin{array}{ll} g_{nm} \longrightarrow h_n^{\alpha \dot{\alpha}} \longrightarrow \{h_n^{\alpha \dot{\alpha}}, \omega_n^{\alpha \beta}, \bar{\omega}_n^{\dot{\alpha} \dot{\beta}}\} & \alpha, \dot{\alpha} = 1, 2 \\ \text{admits a natural generalization to} & s \geq 2 \\ \varphi_{n_1 \dots n_s} \rightarrow h_n^{\alpha_1 \dots \alpha_{s-1} \dot{\alpha}_1 \dots \dot{\alpha}_{s-1}} \rightarrow & \omega_{\alpha_1 \dots \alpha_n, \dot{\beta}_1 \dots \dot{\beta}_m}, & n+m = 2(s-1) \\ s = 1: & \omega(x) = dx^n \omega_n(x) \\ s = 2: & \omega_{\alpha \dot{\beta}}(x), & \omega_{\alpha \beta}(x), & \bar{\omega}_{\dot{\alpha} \dot{\beta}}(x) \\ s = 3/2: & \omega_{\alpha}(x), & \bar{\omega}_{\dot{\alpha}}(x) \end{array}$ Frame-like fields: |n-m| = 0 (bosons) or |n-m| = 1 fermions

By virtue of constraints  $t = \left[\frac{1}{2}|n-m|\right]$  is an order of derivatives  $\omega_{\alpha_1...\alpha_n,\dot{\alpha}_1...\dot{\alpha}_m} = \Pi\left(\partial^t h_{\alpha_1...\alpha_{n_0}\dot{\alpha}_1...\dot{\alpha}_{m_0}}\right), \qquad |n_0-m_0| < 2$ 

#### First-Order Unfolded Equations: Spin Two

**Fields:**  $h^{\alpha \dot{\alpha}}$  and  $\omega^{\alpha \beta}$ ,  $\bar{\omega}^{\dot{\alpha} \dot{\beta}}$ . **Zero-torsion and Einstein equations** 

$$R_{\alpha\dot{\alpha}} = 0, \qquad R_{\alpha_{1}\alpha_{2}} = h^{\alpha_{3}}{}_{\dot{\alpha}} \wedge h^{\alpha_{4}\dot{\alpha}}C_{\alpha_{1}...\alpha_{4}}, \qquad \overline{R}_{\dot{\alpha}_{1}\dot{\alpha}_{2}} = h_{\alpha}{}^{\dot{\alpha}_{3}} \wedge h^{\alpha\dot{\alpha}_{4}}\overline{C}_{\dot{\alpha}_{1}...\dot{\alpha}_{4}}$$

$$C_{\alpha_{1}...\alpha_{4}} \text{ and } \overline{C}_{\dot{\alpha}_{1}...\dot{\alpha}_{4}} : \text{ Weyl tensor}$$
Bianchi identities + Einstein equations imply
$$D^{L}C_{\alpha_{1}...\alpha_{n+4},\dot{\alpha}_{1}...\dot{\alpha}_{n}} + h^{\alpha_{n+5}\dot{\alpha}_{n+1}}C_{\alpha_{1}...\alpha_{n+5},\dot{\alpha}_{1}...\dot{\alpha}_{n+1}} = O(C^{2})$$

$$D^{L}\overline{C}_{\alpha_{1}\dots\alpha_{n},\dot{\alpha}_{1}\dots\dot{\alpha}_{n+4}} + h^{\alpha_{n+1}\dot{\alpha}_{n+5}}\overline{C}_{\alpha_{1}\dots\alpha_{n+1},\dot{\alpha}_{1}\dots\dot{\alpha}_{n+5}} = O(C^{2})$$

 $C_{\alpha_1...\alpha_{n+4},\dot{\alpha}_1...\dot{\alpha}_n}$  and  $\overline{C}_{\alpha_1...\alpha_n,\dot{\alpha}_1...\dot{\alpha}_{n+4}}$ : order-*n* on-shell nontrivial derivatives of the Weyl tensor

Analogously, for general-spin 0-forms

 $C_{\alpha_1\dots\alpha_n\,,\dot{\beta}_1\dots\dot{\beta}_m}\,,\qquad |n-m|=2s$ 

$$s = 0: \quad C(x)$$

$$s = 1/2: \quad C_{\alpha}(x), \qquad \bar{C}_{\dot{\alpha}}(x)$$

$$s = 1: \quad C_{\alpha\beta}, \qquad \bar{C}_{\dot{\alpha}\dot{\beta}}$$

$$s = 3/2: \quad C_{\alpha\beta\gamma}, \qquad \bar{C}_{\dot{\alpha}\dot{\beta}\dot{\gamma}}$$

$$s = 2: \quad C_{\alpha_1...\alpha_4}, \qquad \bar{C}_{\dot{\alpha}_1...\dot{\alpha}_4}$$

#### **Central On-Shell Theorem**

#### Infinite set of integer spins

 $\omega(y,\bar{y} \mid x), \quad C(y,\bar{y} \mid x) \quad f(y,\bar{y}) = \sum_{n,m=0}^{\infty} \frac{1}{n!m!} f_{\alpha_1...\alpha_n,\dot{\alpha}_1...\dot{\alpha}_m} y^{\alpha_1} \dots y^{\alpha_n} \bar{y}^{\dot{\alpha}_1} \dots \bar{y}^{\dot{\alpha}_n}$ The full unfolded system for free bosonic fields is 1989

$$\star \qquad R_1(y,\overline{y} \mid x) = \frac{i}{4} \left( \eta \overline{H}^{\dot{\alpha}\dot{\beta}} \frac{\partial^2}{\partial \overline{y}^{\dot{\alpha}} \partial \overline{y}^{\dot{\beta}}} C(0,\overline{y} \mid x) + \overline{\eta} H^{\alpha\beta} \frac{\partial^2}{\partial y^{\alpha} \partial y^{\beta}} C(y,0 \mid x) \right)$$
  
$$\star \star \qquad \tilde{D}_0 C(y,\overline{y} \mid x) = 0$$

**Vacuum:**  $sp(4) \sim o(3, 2)$ 

$$\begin{split} \mathbf{R}_{\alpha\beta} &:= \mathbf{d}\omega_{\alpha\beta} + \omega_{\alpha\gamma}\omega_{\beta}{}^{\gamma} - \mathbf{H}_{\alpha\beta} = \mathbf{0} \,, \qquad \mathbf{R}_{\alpha\dot{\beta}} := \mathbf{d} + \omega_{\alpha\gamma}\mathbf{h}^{\gamma}{}_{\dot{\beta}} + \overline{\omega}_{\dot{\beta}\dot{\delta}}\mathbf{h}_{\alpha}{}^{\delta} = \mathbf{0} \\ \mathbf{H}^{\alpha\beta} &:= \mathbf{h}^{\alpha\dot{\alpha}} \wedge \mathbf{h}^{\beta}{}_{\dot{\alpha}} \,, \qquad \overline{\mathbf{H}}^{\dot{\alpha}\dot{\beta}} := \mathbf{h}^{\alpha\dot{\alpha}} \wedge \mathbf{h}_{\alpha}{}^{\dot{\beta}} \\ R_{1}(y, \overline{y} \mid x) = D_{0}^{ad}\omega(y, \overline{y} \mid x) \qquad D_{0}^{ad} = D^{L} - h^{\alpha\dot{\beta}} \Big( y_{\alpha}\frac{\partial}{\partial \overline{y}^{\dot{\beta}}} + \frac{\partial}{\partial y^{\alpha}}\overline{y}_{\dot{\beta}} \Big) \\ \widetilde{D}_{0} &= D^{L} + h^{\alpha\dot{\beta}} \Big( y_{\alpha}\overline{y}_{\dot{\beta}} + \frac{\partial^{2}}{\partial y^{\alpha}\partial \overline{y}^{\dot{\beta}}} \Big) \qquad D^{L} = \mathsf{d}_{x} - \Big( \omega^{\alpha\beta}y_{\alpha}\frac{\partial}{\partial y^{\beta}} + \overline{\omega}^{\dot{\alpha}\dot{\beta}}\overline{y}_{\dot{\alpha}}\frac{\partial}{\partial \overline{y}^{\dot{\beta}}} \Big) \end{split}$$

**\*\*** implies that higher-order terms in y and  $\bar{y}$  describe higher-derivative descendants of the primary HS fields

### **Fields of the Nonlinear System**

Closed formulation of nonlinear equations demands the doubling of spinors and Klein operator

$$\omega(Y|x) \longrightarrow W(Z;Y;k|x), \qquad C(Y|x) \longrightarrow B(Z;Y;k|x)$$

Some of the nonlinear HS equations determine the dependence on the additional variables  $Z_A$  in terms of "initial data"  $\omega(Y;k|x) := W^{dyn}(0;Y|x) + W^{top}(0;Y|x)k$  $C(Y;k|x) := B^{dyn}(0;Y|x)k + B^{top}(0;Y|x)$  $S(Z;Y;k|x) = dZ^A S_A(Z;Y;k|x)$  is a connection along  $Z^A$ Topological fields: finite # d.o.f.: tensors

Klein operator k generates chirality automorphisms

P

$$kf(A) = f(\tilde{A})k, \quad A = (a_{\alpha}, \bar{a}_{\dot{\alpha}}) : \quad \tilde{A} = (-a_{\alpha}, \bar{a}_{\dot{\alpha}})$$
$$(Y) = P^{\alpha \dot{\alpha}} y_{\alpha} \bar{y}_{\dot{\alpha}} \longrightarrow \quad \tilde{P}(Y) = -P(Y), \qquad \tilde{M}(Y) = M(Y)$$

### **Nonlinear HS Equations**

**HS** star product

$$(f \star g)(Z, Y) = \int dS dT \exp iS_A T^A f(Z + S, Y + S)g(Z - T, Y + T)$$

 $[Y_A, Y_B]_{\star} = -[Z_A, Z_B]_{\star} = 2iC_{AB}, \qquad \qquad Z - Y : Z + Y \text{ normal ordering}$ 

#### **Inner Klein operators:**

$$\kappa = \exp i z_{\alpha} y^{\alpha}, \qquad \bar{\kappa} = \exp i \bar{z}_{\dot{\alpha}} \bar{y}^{\dot{\alpha}}, \qquad \kappa \star f = \tilde{f} \star \kappa, \qquad \kappa \star \kappa = 1$$

$$\begin{cases} dW + W \star W = 0 \\ dB + W \star B - B \star W = 0 \\ dS + W \star S + S \star W = 0 \end{cases}$$

$$g \star B - B \star S = 0 \\ S \star B - B \star S = 0 \\ S \star S = \mathbf{i} (\mathbf{dZ}^{\mathbf{A}} \mathbf{dZ}_{\mathbf{A}} + \eta \mathbf{dz}^{\alpha} \mathbf{dz}_{\alpha} \mathbf{B} \star \mathbf{k} \star \kappa + \bar{\eta} \mathbf{d} \bar{z}^{\dot{\alpha}} \mathbf{d} \bar{z}_{\dot{\alpha}} \mathbf{B} \star \mathbf{k} \star \bar{\kappa})$$

Real physics is localized in the x-independent twistor sector

### **Perturbative Analysis**

#### **Vacuum solution**

$$B_0 = 0$$
,  $S_0 = dZ^A Z_A$ ,  $W_0 = \frac{1}{2} \omega_0^{AB}(x) Y_A Y_B$   
 $dW_0 + W_0 \star W_0 = 0$ 

 $\omega_0^{AB}(x)$ : describes  $AdS_4$ .

**First-order fluctuations** 

 $B_1 = C(Y), \qquad S = S_0 + S_1, \qquad W = W_0(Y) + W_1(Y) + W_0(Y)C(Y)$ 

$$[S_0, f]_{\star} = -2i \mathsf{d}_Z f, \qquad \mathsf{d}_Z = dZ^A \frac{\partial}{\partial Z^A}$$

#### **Reconstruction of** $Z^A$ Variables

#### Perturbatively, equations containing S have the form

 $\mathsf{d}_Z U_n(Z;Y|dZ) = V[U_{< n}](Z;Y|dZ) \qquad \mathsf{d}_Z V[U_{< n}](Z;Y|dZ) = 0$ 

can be solved as

 $U_n(Z;Y|dZ) = \mathsf{d}_Z^* V[U_{< n}](Z;Y|dZ) + \mathbf{h}(\mathbf{Y}) + \mathsf{d}_Z \epsilon(Z;Y|dZ)$ 

For instance

$$d_Z^*V(Z;Y|dZ) = Z^A \frac{\partial}{dZ^A} \int_0^1 \frac{dt}{t} V(tZ;Y|tdZ)$$

Alternative  $d_Z^*$  that differ by  $d_Z$ -closed forms can also be used. Proper choice of boundary conditions in Z-variables is most important in the context of locality beyond the free field level!

Nontrivial space-time equations on  $\omega(Y|x)$  and C(Y|x) are in the sector of d<sub>Z</sub>-cohomology

Central On-Shell Theorem is reproduced in the lowest order

#### **Conserved Currents and Current Deformation**

Gauge invariant conserved currents  $J(Y_1, Y_2|x)$  are represented by the bilinears of C(Y|x)

$$J(Y_1, Y_2|x) := C(Y_1|x)\tilde{C}(Y_2|x), \qquad \tilde{C}(y, \bar{y}|x) = C(-y, \bar{y}|x)$$

As a consequence of the rank-one equation for C(Y|x),  $J(Y_1, Y_2|x)$  obeys the current equation Gelfond, MV (2003)

$$\tilde{D}_2 J(Y_1, Y_2 | x) = 0, \qquad \tilde{D}_2 := D^L - i\lambda h^{\alpha \dot{\beta}} \Big( y_{1\alpha} \bar{y}_{1\dot{\beta}} - y_{2\alpha} \bar{y}_{2\dot{\beta}} - \frac{\partial^2}{\partial y_1^{\alpha} \partial \bar{y}_1^{\dot{\beta}}} + \frac{\partial^2}{\partial y_2^{\alpha} \partial \bar{y}_2^{\dot{\beta}}} \Big)$$

#### Current deformation has a form of a linear system

$$D\omega - L(w,C) + \Gamma(w,J) = 0,$$

$$\tilde{D}C + \mathcal{H}(w, J) = 0, \qquad \tilde{D}_2 J(Y_1, Y_2 | x) = 0$$

$$L(w,C) := \frac{i}{4} \left( \eta \overline{H}^{\dot{\alpha}\dot{\beta}} \frac{\partial^2}{\partial \overline{y}^{\dot{\alpha}} \partial \overline{y}^{\dot{\beta}}} \ \overline{C}(0,\overline{y}|x) + \overline{\eta} H^{\alpha\beta} \frac{\partial^2}{\partial y^{\alpha} \partial y^{\beta}} \ C(y,0|x) \right)$$

Linear functionals  $\Gamma$  and  $\mathcal{H}$  obey the compatibility conditions

### Locality and Nonlocality in HS Gauge Theory

Having infinitely many HS fields with higher derivatives in interactions, the HS Gauge Theory is not local

$$\lambda^{-1}D \sim 1$$
 since  $[\lambda^{-1}D, \lambda^{-1}D] \sim 1$ 

A different mass parameter like  $\alpha'$  is needed for a low-energy expansion

In HS equations, nonlocality is due to infinite tails of contractions

$$\int \frac{d^4 S d^4 T}{(2\pi)^4} \exp i[s_\beta t^\beta + \bar{s}_{\dot{\beta}} \bar{t}^{\dot{\beta}}] J(y+s,\bar{y}+\bar{s};y+t,\bar{y}+\bar{t})$$
  
=  $\exp -i[\partial_{1\alpha}\partial_{2\beta}\epsilon^{\alpha\beta} + \bar{\partial}_{1\dot{\alpha}}\bar{\partial}_{2\dot{\beta}}\epsilon^{\dot{\alpha}\dot{\beta}}] J(y,\bar{y};y,\bar{y})$ 

# of derivatives in y and  $\bar{y}$  infinitely increases for given helicities  $\Rightarrow$ infinite tails of space-time derivatives and hence nonlocality True locality: absence of integration over s and t or  $\bar{s}$  and  $\bar{t}$ 

$$\int \frac{d\bar{s}d\bar{t}}{(2\pi)^2} \exp i[\bar{s}_{\dot{\beta}}\bar{t}^{\dot{\beta}}]f(y,\bar{y}+\bar{s})g(y,\bar{y}+\bar{t}) = f(y,\bar{y}) \exp[-i\overleftarrow{\partial}_{\dot{\alpha}}\overrightarrow{\partial}_{\dot{\beta}}\epsilon^{\dot{\alpha}\dot{\beta}}]g(y,\bar{y})$$

For given helicities carried by g and f, only a single term in the sum contributes hence containing a finite number of derivatives

## **Resolution Redefinition**

#### Original homotopy operator gives in the $\eta$ -sector

$$B_{2\eta} = \frac{\eta}{2} \int d_{+}^{4} \tau \delta (1 - \sum_{i=1}^{4} \tau_{i}) y^{\beta} \left( \delta(\tau_{1}) \partial_{2\beta} + \delta(\tau_{2}) \partial_{1\beta} \right) \exp(X) C(Y_{1}; K) C(Y_{2}; K) * k * \kappa |_{Y_{1}}$$

 $X = i(1-\tau_3)z_{\alpha}y^{\alpha} + \tau_3y^{\alpha}(\partial_{1\alpha} + \partial_{2\alpha}) + z^{\alpha}(\tau_2\partial_{2\alpha} - \tau_1\partial_{1\alpha}) + i(\tau_3 + \tau_4)\partial_{1\alpha}\partial_2^{\alpha} + i\bar{\partial}_{1\dot{\alpha}}\bar{\partial}_2^{\dot{\alpha}},$ 

$$\partial_{i\alpha} := \frac{\partial}{\partial y_i^{\alpha}}, \qquad \bar{\partial}_{i\dot{\alpha}} := \frac{\partial}{\partial \bar{y}_i^{\dot{\alpha}}}$$

Since  $\tau_4 \ge 0$ , the expansion coefficients  $\tau_3 + \tau_4$  in powers of  $\partial_{1\alpha} \partial_2^{\alpha}$  are larger than those  $(\tau_3)$  of  $y^{\alpha} \partial_{i\alpha}$ .

#### Using the Schoutens identity

$$z_{\alpha}y^{\alpha}\partial_{1\beta}\partial_{2}^{\beta} + z^{\alpha}\partial_{1\alpha}y^{\beta}\partial_{2\beta} - z^{\alpha}\partial_{2\alpha}y^{\beta}\partial_{1\beta} = 0$$

redefinition of the resolution operator

$$\mathsf{d}_{loc}^*B := \mathsf{d}_Z^*B - \Delta C_{2\eta}$$

## **Final result**

$$B_{2\eta}^{loc} = \frac{1}{2} \eta \int d_{+}^{3} \tau \left( \delta' (1 - \sum_{i=1}^{3} \tau_{i}) - iy_{\alpha} z^{\alpha} \delta (1 - \sum_{i=1}^{3} \tau_{i}) \right)$$
$$\int d^{4} U d^{4} V \exp i (U_{A} V^{A} + (1 - \tau_{3}) z_{\alpha} y^{\alpha})$$
$$C(\tau_{3} y - \tau_{1} z + \tau_{3} u, \bar{y} + \bar{u}; K) C(\tau_{3} y + \tau_{2} z + v, \bar{y} + \bar{v}; K) * k * \kappa.$$

The coefficient responsible for the index contraction between the first and second factors of C equals to those in front of the y variables, analogously to the star product of Z-independent functions.

 $\Delta C_{2\eta}$  represents the nonlinear shift of 1605.02662 reducing the nonlocal bilinear corrections to the local form in the sector of x, y-variables.

$$\tilde{D}C + \frac{\eta}{4} \int \frac{d\bar{s}d\bar{t}}{(2\pi)^2} \exp i[\bar{s}_{\dot{\beta}}\bar{t}^{\dot{\beta}}] \int_0^1 d\tau h^{\alpha\dot{\alpha}} y_\alpha(\tau\bar{s}_{\dot{\alpha}} + (1-\tau)\bar{t}_{\dot{\alpha}}) \\J(\tau y, (\tau-1)y; \bar{y} + \bar{s}, \bar{y} + \bar{t}; K) * k = 0$$

## **Current Contribution to the Gauge Sector**

**Contribution to spin-s equations** 

Gelfond, MV (2017)

$$\mathrm{d}\omega + \widetilde{\mathcal{J}}_{\mathrm{ss}} + \ldots = 0 \, ,$$

$$\begin{aligned} \widetilde{\mathcal{J}}_{ss} &= i\eta \bar{\eta} \frac{(s-2)!}{8(2s)!} \sum_{k,m \in [0,s]} \frac{(m+k)!(2s-m-k)!}{(s-k)!k!(s-m)!m!} \left( \mathcal{N}_1 \right)^m \left( -\mathcal{N}_2 \right)^{s-m} \left( -\overline{\mathcal{N}}_2 \right)^k \left( \overline{\mathcal{N}}_1 \right)^{s-k} \\ &\left\{ \sum_{0 \le n \le s} \frac{1}{(s+n-1)!} (i\partial_1 \gamma \partial_2 \gamma)^n \sum_{j,l=0,1} C^{j,1-j} (Y^1|x) k^j \bar{k}^{1-j} C^{l,1-l} (Y^2|x) k^l \bar{k}^{1-l} \right. \\ &\left. + \sum_{0 < n \le s} \frac{1}{(s+n-1)!} (i\bar{\partial}_1 \gamma \bar{\partial}_2 \gamma)^n \sum_{j,l=0,1} C^{j,1-j} (Y^1|x) k^j \bar{k}^{1-j} C^{l,1-l} (Y^2|x) k^l \bar{k}^{1-l} \right\} \right|_{Y^j=0} \\ &\left. \mathcal{N}_j = y^\alpha \partial_{j\alpha} \,, \qquad \overline{\mathcal{N}}_j = \bar{y}^{\dot{\alpha}} \bar{\partial}_{j\dot{\alpha}} \end{aligned}$$

Although  $\widetilde{\mathcal{J}}_{s,s} \sim \eta \overline{\eta}$  the current contribution to Fronsdal equtions depends on the phase of  $\eta$  through Central on-shell theorem. Matching with the vertex obtained by Metsaev in 1991 from the higher-order analysis and by Taronna and Sleight in 2016 from the

**holographic analysis at**  $\eta = 1$  Misuna (2017)

# **HS** holography

General idea of HS dualitySundborg (2001), Witten (2001) $AdS_4/CFT3$ :Klebanov, Polyakov (2002); Giombi, Yin (2009,2010)

Extension to duality between CS boundary theory and HS theories in  $AdS_4$  for various  $\eta = \exp i \frac{\pi}{2} \lambda_b$  with  $\lambda_b = \frac{N}{k}$ Aharony, Gur-Ari, Yacoby (2011); Giombi, Minwalla, Prakash, Trivedi, Wadia, Yin Explicit check was only partially successful Giombi and Yin (2012)

Local frame allows further HS holography checks: zero-form sector Sezgin, Skvortsov, Zhu (May 2017), Didenko, MV (May 2017) one-form sector N.Misuna (June 2017)

## **Functional Class**

Appropriate space  $V_{0,0}$  is spanned (2015) by functions of the form

$$f(Z;Y) = \int_0^1 d\tau \phi(\tau Z; (1-\tau)Y;\tau) \exp i\tau Z_A Y^A$$

with  $\phi(W; U; \tau)$  regular in W and U and integrable in  $\tau$ .

Analysis of locality in the zero-form sector clarifies the structure of the pre-exponential factor on  $\tau_i$  in the nonlinear case

$$\int d_{+}^{5} \tau \rho(\tau) \int d^{2}u d^{2}v \exp i(u_{\alpha}v^{\alpha} + tz_{\alpha}y^{\alpha})C(\tau_{1}y - \tau_{3}z + \tau_{5}u)C(-\tau_{4}z - \tau_{2}y + v) * k$$
  
$$\tau_{3} \leq t , \qquad \tau_{4} \leq t , \qquad \tau_{1} \leq 1 - t , \qquad \tau_{2} \leq 1 - t , \qquad \tau_{5} \leq t$$

The restriction on  $\tau_5$  extends the 2015 results to the nonlinear case.

Such functions induce minimally nonlocal higher-order corrections! Special form of the proper nonlinear correction is not visible from the analysis of quadratic corrections at Z = 0!

### Conclusion

To get local (or minimally nonlocal) results at higher orders one has to use a particular resolution operator  $d_{loc}^*$ .

HS equations in the local frame reproduce correctly local corrections and anticipated AdS/CFT results: Gelfond, Didenko, Misuna (2017)

Classes of functions in the twistor space distinguishing between local and nonlocal field redefinitions in HS theory are identified

## To do

Further computations in standard HS holography for  $\forall \eta$ Via one-form sector On-shell invariant functionals MV 2015

Analysis of  $d_{loc}^*$ , (non?)locality and HS holography at higher orders