Non-relativistic Limits	Non-relativistic String theory	Non-relativistic NS-NS Gravity	Minimal Supergravity	Outlook
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Advances in Non-relativistic Quantum Gravity

Eric Bergshoeff

Groningen University

work done in collaboration with

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International Conference on Strings, Fields and Holograms

Ascona, October 11 2021

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Three Roads to Quantum Gravity



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NR Quantum Gravity

Does combining gravity with quantum mechanics require relativity?

Does NR string theory define NR quantum gravity?

Does NR gravity has its own holographic principle?

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Some References

NR gravity/string theory involving null-reduction

T. Harmark, J. Hartong and N. A. Obers (2017); Kluson (2018);

T. Harmark, J. Hartong, L. Menculini, N. A. Obers and Z. Yan (2018);

Kluson (2019); Roychowdhury (2019); T. Harmark, J. Hartong, L. Menculini,

N. A. Obers and G. Oling (2019); A.D. Gallegos, U. Gürsoy and N. Zinnato (2019);

L. Bidussi, T. Harmark, J. Hartong, N.A, Obers, G. Oling (2021)

NR strings with NR worldsheet

C. Batlle, J. Gomis and D. Not (2017); C. Batlle, J. Gomis, L. Mezincescu and P. K. Townsend (2017); T. Harmark, J. Hartong and N. A. Obers (2017); T. Harmark, J. Hartong, L. Menculini, N. A. Obers and Z. Yan (2018)

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Non-relativistic String theory



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Non-relativistic NS-NS Gravity



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Defining a NR Limit

STEP 1: decomposing $E_{\mu}{}^{\hat{A}} = (E_{\mu}{}^{0}, E_{\mu}{}^{A'}) = (\text{clock, ruler})$ and introducing M_{μ} , perform an invertable field redefinition involving a parameter c:

$$E_{\mu}{}^{0} = c au_{\mu} + c^{-1} m_{\mu} \,, \quad E_{\mu}{}^{A'} = e_{\mu}{}^{A'} \,, \quad M_{\mu} = c au_{\mu} - c^{-1} m_{\mu}$$

STEP 2: take the limit $c \rightarrow \infty$ and take care of possible divergences

Example: Particles and a 'critical' limit

cp. to Seiberg, Susskind, Toumbas (2000); Gopakumar, Maldacena, Minwalla, Strominger (2000);

Danielsson, Guijosa, Kruczenski (2000), Gomis, Ooguri (2001)

Starting from a particle coupled to gravity, the red terms in the above field redefinition lead to divergencies in the kinetic and Wess-Zumino term that cancel against each other.



Infinities

Using a second-order formulation of general relativity, the NR limit of the spin-connection fields $\Omega_{\mu}{}^{\hat{A}\hat{B}}(E)$ contains a leading divergence that usually is set to zero by imposing the zero torsion constraint

$\partial_{[\mu}\tau_{\nu]}=0$

Given this constraint the NR limit of the Einstein e.o.m. (no action!) yields the NC gravity e.o.m. where the Newton potential Φ can be identified as the time component of the central charge gauge field m_{μ} :

$$\Phi \sim au^{\mu} m_{\mu}$$

This NC gravity theory is a reformulation of Newtonian gravity valid in any frame and including strong gravity effects

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The Zero Torsion Constraint

$$\partial_{[\mu} \tau_{\nu]} = 0 \quad \rightarrow \quad \tau_{\mu} = \partial_{\mu} \rho \quad \text{with} \quad \tau_{\mu} \quad \text{clock function}$$



$$\Delta T = \int_{\mathcal{C}} \mathrm{d}x^{\mu} \tau_{\mu} = \int_{\mathcal{C}} \mathrm{d}
ho \, \text{ is path-independent } \,
ightarrow \, ext{absolute time}$$

Torsional NC gravity : $\partial_{\mu}\tau_{\nu} - \Gamma_{\mu\nu}^{\ \rho}\tau_{\rho} = 0 \quad \rightarrow \quad \Gamma_{[\mu\nu]}^{\ \rho}\tau_{\rho} = \partial_{[\mu}\tau_{\nu]}$

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Geometry with Co-dimension 2 Foliation

The string should be coupled to a 2-form gauge field $B_{\mu\nu}$ with

$$B_{\mu\nu} = -c^2 \epsilon_{AB} \tau_{\mu}{}^A \tau_{\nu}{}^B + b_{\mu\nu}$$

defining a geometry with a co-dimension 2 foliation where $\tau_{\mu} \rightarrow \tau_{\mu}{}^{A}$ with $\hat{A} = (A, A') = (0, 1, A')$



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The Basic Variables

The decomposition leading to NC gravity

$$\{E_{\mu}^{\hat{A}}, M_{\mu}\} \rightarrow \{\tau_{\mu}, e_{\mu}^{A'}, m_{\mu}\}$$

gets replaced by the following redefinition:

$$\{E_{\mu}{}^{\hat{A}}, B_{\mu\nu}, \Phi\} \rightarrow \{\tau_{\mu}{}^{A}, e_{\mu}{}^{A'}, b_{\mu\nu}, \phi\}$$

The Newton potential Φ can be identified with the time-space component $\epsilon^{AB} \tau^{\mu}{}_{A} \tau^{\nu}{}_{B} b_{\mu\nu}$ of the 2-form gauge field $b_{\mu\nu}$

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The NR String Sigma Model

The bosonic closed string sigma model (without Yang-Mills) is given by

J. Gomis, Z. Yan + E.B. (2018); J. Gomis, J. Rosseel, C. Şimşek, Z. Yan + E.B. (2019)

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$$S_{\rm NR\sigma} = -\frac{T}{2} \int d^2 \sigma \left[\sqrt{-h} \, h^{\alpha\beta} \, \partial_\alpha x^\mu \partial_\beta x^\nu e_\mu^{\ A'} e_\nu^{\ B'} \delta_{A'B'} + \, \epsilon^{\alpha\beta} \partial_\alpha x^\mu \partial_\beta x^\nu b_{\mu\nu} \right] + S_{\rm dilaton}$$

with world-sheet metric
$$h_{\alpha\beta} \sim \tau_{\alpha\beta} \equiv \partial_{\alpha} x^{\mu} \partial_{\beta} x^{\nu} \tau_{\mu}^{\ A} \tau_{\nu}^{\ B} \eta_{AB}$$

This is the generalization of flat spacetime to a string NC background Gomis, Ooguri (2001); Danielsson, Guijosa, Kruczenski (2000)

Note: we have not imposed any geometric constraint sofar (they follow later from dynamics)



Special Features

• The KR 2-form field $b_{\mu\nu}$ transforms under string-Galilean boost transformations:

$$\delta b_{\mu\nu} = \partial_{[\mu} \lambda_{\nu]} + 2\epsilon_{AB} \lambda_{A'}{}^{A} \tau_{[\mu}{}^{B} e_{\nu]}{}^{A'}$$

A relativistic matter field $B_{\mu\nu}$ becomes a NR geometric field $b_{\mu\nu}$

• There is an emergent dilatation symmetry:

$$\delta \tau_{\mu}{}^{A} = \lambda_{D} \tau_{\mu}{}^{A}, \qquad \qquad \delta \phi = \lambda_{D}$$

This means that the # of NR background fields is one less than the # of relativistic background fields \rightarrow one 'missing e.o.m.'

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Canceling the Divergences

$$S_{\rm rel} = rac{1}{2\kappa^2} \int d^{10} x \, E\!\left(\mathcal{R} - rac{1}{12} \mathcal{H}_{\mu
u
ho} \mathcal{H}^{\mu
u
ho}
ight)$$

with $\mathcal{H}_{\mu\nu\rho} = 3\partial_{[\mu}B_{\nu\rho]}$. We redefine

$$E_{\mu}{}^{A} = c \tau_{\mu}{}^{A}, \qquad E_{\mu}{}^{A'} = e_{\mu}{}^{A'}, \qquad B_{\mu\nu} = -c^{2} \epsilon_{AB} \tau_{\mu}{}^{A} \tau_{\nu}{}^{B} + b_{\mu\nu}$$

and find

$$S = c^{2} \frac{\binom{2}{5}}{5} + \frac{\binom{0}{5}}{5} + c^{-2} \frac{\binom{-2}{5}}{5} + c^{-4} \frac{\binom{-4}{5}}{5}$$

where $\stackrel{(2)}{S}$ is proportional to the torsion tensor

$$\tau_{\mu\nu}{}^{A} \equiv \partial_{[\mu}\tau_{\nu]}{}^{A}$$

Two miracles: (i) the metric and 2-form contributions to $\stackrel{(2)}{S}$ precisely cancel and (ii) $\stackrel{(0)}{S}$ is invariant under local dilatations!

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Special Features of Non-relativistic Action

$$\begin{split} S_{\rm NR} &= \frac{1}{2 \,\kappa^2} \int {\rm d}^{10} x \, e \left({\rm R}(J) - \frac{1}{12} \, h_{A'B'C'} h^{A'B'C'} \right. \\ & - 4 \, \mathcal{D}_{A'} b^{A'} - 4 \, b_{A'} b^{A'} - 4 \, \tau_{A'\{AB\}} \tau^{A'\{AB\}} \right). \end{split}$$

- the action has an emergent dilatation symmetry and therefore has one 'missing field' and one 'missing e.o.m.'
 - the dilatation gauge field $b_{\mu}=b_{\mu}(au,\phi)$ is dependent
- The 'missing' e.o.m. follows from taking the NR limit of the e.o.m. and is precisely the Poisson equation of the Newton potential
 - The full set of e.o.m. form a reducible, but indecomposable representation: the Poisson equation needs the action!
- The e.o.m. of the Newton potential itself gives the following non-linear geometric constraint: $\tau_{B'C'A}\tau^{B'C'A} = 0$
 - This prevents an overdetermined set of equations

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Action, E.O.M. and β -Functions

The e.o.m. of non-relativistic string theory are determined by calculating the β -functions

Gomis, Oh, Yan (2019), Yan, Yu (2019), Gomis, Yan, Yu (2020); see also Gallegos, Gürsoy and Zinnatos (2019)

The emergent dilatation symmetry has the following effect:

NR NS-NS action	\rightarrow	common equations	+ Non-linear
NR $\beta\text{-functions}$	\rightarrow	common equations	+ Poisson
NR e.o.m.	\rightarrow	common equations	+ Poisson + Non-linear

The nonlinear equation is required in order that the NR string σ model does not flow towards a relativistic string σ model

Z. Yan (2021)

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Torsional String Newton-Cartan (TSNC) geometry

basic variables:
$$\{\tau_{\mu}{}^{A}, e_{\mu}{}^{A'}, b_{\mu\nu}, \phi\}$$

 $\{\omega_{\mu}, \omega_{\mu}{}^{AA'}, \omega_{\mu}{}^{A'B'}, b_{\mu}\} \text{ are dependent, e.g., } b_{\mu} = e_{\mu}{}^{A'} \tau_{A'A}{}^{A} + \tau_{\mu}{}^{A} \partial_{A} \phi$

$$\nabla_{\mu}\tau_{\nu}{}^{A} \equiv \partial_{\mu}\tau_{\nu}{}^{A} - \omega_{\mu} \epsilon^{AB}\tau_{\nu B} - \mathbf{b}_{\mu}\tau_{\nu}{}^{A} - \Gamma^{\rho}_{\mu\nu}\tau_{\rho}{}^{A} = 0,$$

$$\nabla_{\mu}e_{\nu}{}^{A'} \equiv \partial_{\mu}e_{\nu}{}^{A'} - \omega_{\mu}{}^{A'B'}e_{\nu B'} + \omega_{\mu}{}^{AA'}\tau_{\nu A} - \Gamma^{\rho}_{\mu\nu}e_{\rho}{}^{A'} = 0$$

non-zero torsion: $T^{\rho}_{\mu\nu} = 2\Gamma^{\rho}_{[\mu\nu]} = 2D_{[\mu}(\omega, b)\tau_{\nu]}^{A}\tau_{A}^{\rho}$

cp. to Geracie, Prabhu, Roberts (2015)

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Taking the NR limit of fermions

Define world-sheet chirality projection operators Π_\pm by

$$\Pi_{\pm} = \frac{1}{2}(1 \pm \gamma_0 \gamma_1)$$

and consider projected target space spinors

$$\chi_{\pm} = \Pi_{\pm} \chi$$

Gomis, Kamimura, Townsend (2004)

We then redefine the spinor χ as

$$\chi = c^{1/2} \chi_+ + c^{-1/2} \chi_-$$

such that, after taking $c \to \infty$, the projected spinors transform under Galilean boosts as follows:

$$\delta\chi_{+} = 0$$
 and $\delta\chi_{-} = \frac{1}{2}\lambda^{AA'}\Gamma_{AA'}\chi_{+}$

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New features compared to the bosonic case

- There is no direct connection between a two-dimensional sigma model description and the NR target space effective action
- Taking the naive NR limit leads to divergent terms in the supersymmetry rules

These divergences can be controlled by

- the occurrence of 2 'superconformal' Stueckelberg symmetries beyond dilatations yielding a shortened supergravity multiplet
- imposing by hand the following twistless torsional constraint:

$$T^{-}_{\mu\nu} = 0$$
 or $\tau_{[\mu}^{-}\partial_{\nu}\tau_{\rho]}^{-} = 0$

Christensen, Hartong, Obers, Rollier (2013)

defining a 'self-dual' TSNC geometry (invariant under SUSY!)

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Minimal Supergravity Action versus E.O.M.

- the action has one emergent dilatation and two emergent superconformal symmetries. It therefore has one 'missing' bosonic and two 'missing' fermionic fields plus corresponding 'missing' e.o.m.
- The 'missing' e.o.m. follow from taking the NR limit of the e.o.m. and are precisely the Poisson equation of the Newton potential plus two fermionic partner equations

• The minimal supergravity action is a pseudo action in the sense that it is only invariant under supersymmetry if one uses the twistless torsional constraint <u>after</u> varying the action. Due to this the e.o.m. that follow from the action transform under supersymmetry to the 'missing' e.o.m.: they belong to the same supermultiplet

cp. to Vanhecke, Van Proeyen (2017)

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Outlook

- how general is our limit technique?
 - invertable field redefinition, cancellation of divergences, local dilatation symmetry
- including Yang-Mills to obtain heterotic gravity
 - sigma model anomaly, T-duality: taking a string NR limit followed by spatial reduction is dual to a null-reduction
- open strings

see lectures by Z. Yan at 1st School on NR QFT, Gravity and Geometry

connection to Double Field Theory

Ko, Melby-Thompson, Meyer and Park (2015); Gallegos, Gürsoy, Verma and Zinnato (2020)

extension to IIA/IIB supergravity and M-theory

for bosonic sector of M-theory, see Blair, Gallegos, Zinnato (2021)

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Take-Home Message

Our results pave the way for a target space approach to NR string theory:

(supersymmetric) brane solutions, compactifications, NR holography etc.

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